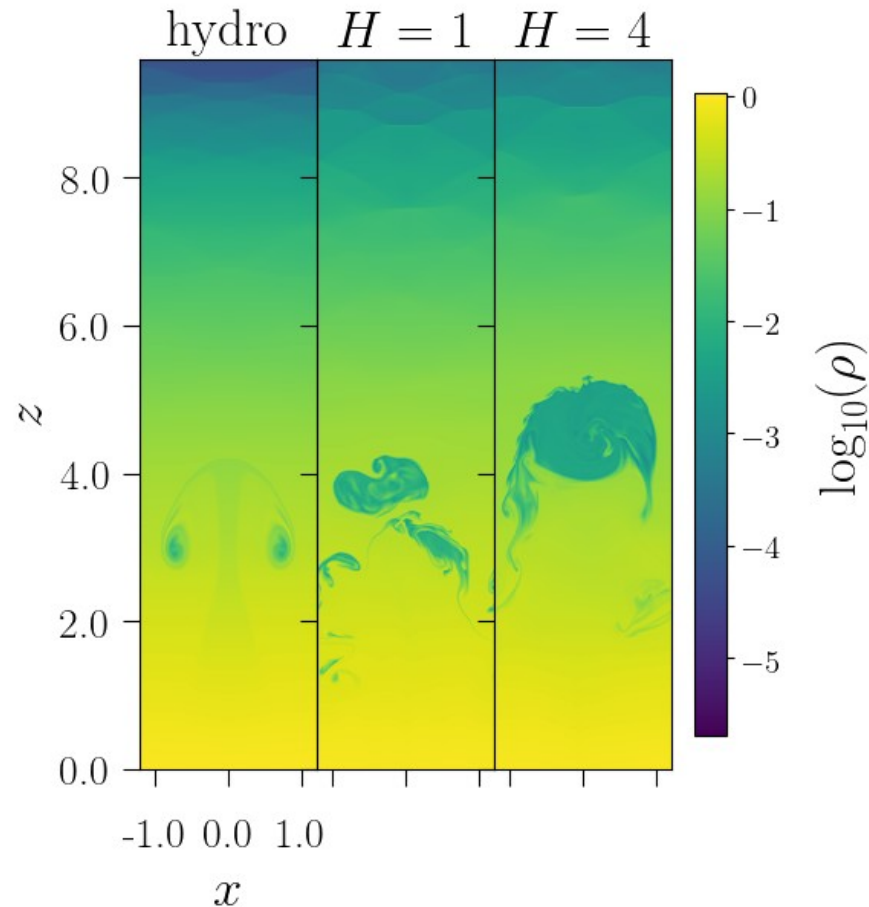
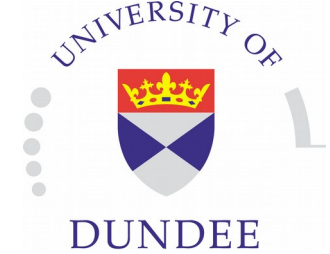
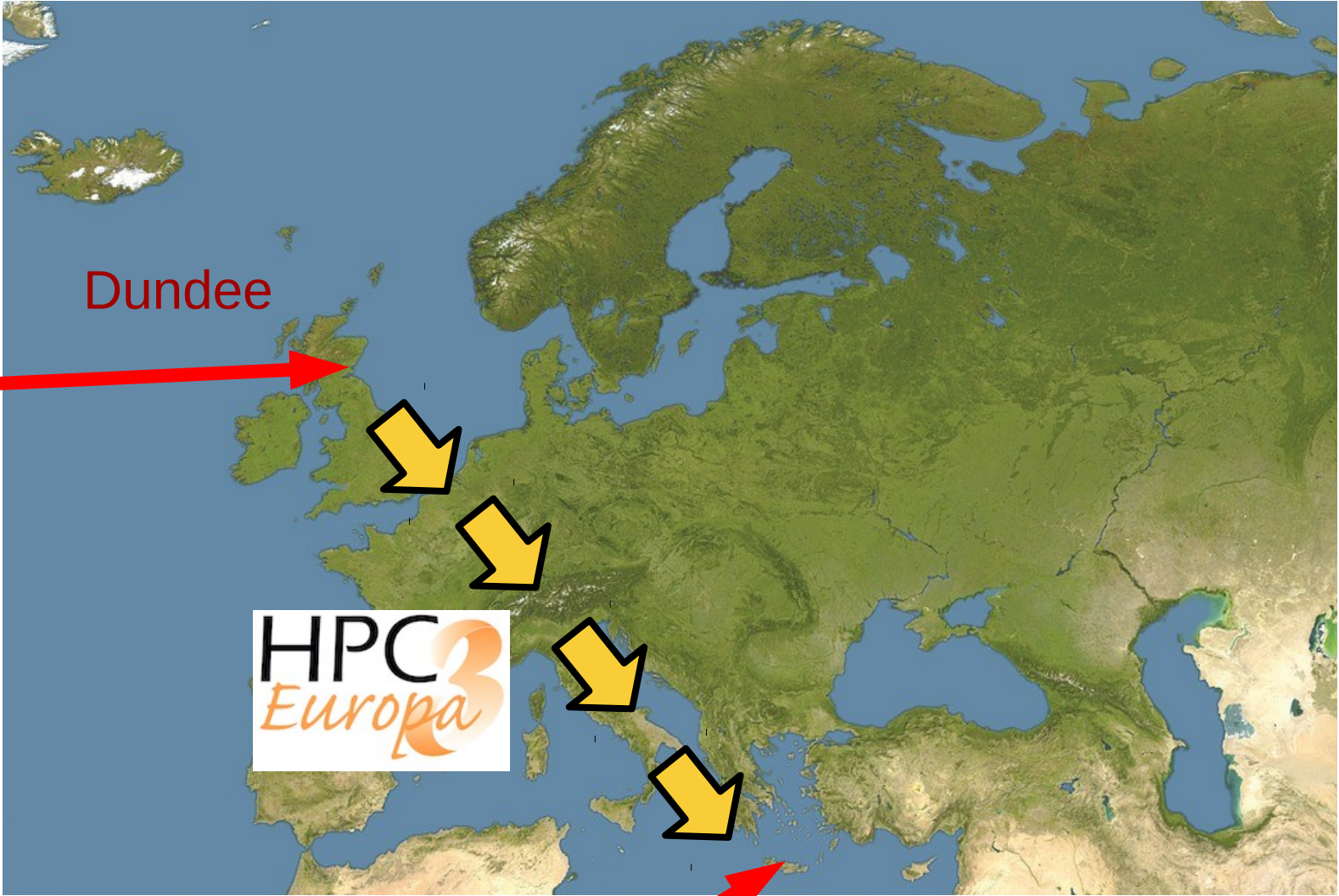


Stabilizing effect of magnetic helicity on magnetic cavities in the intergalactic medium.

Simon Candelaresi, Fabio Del Sordo



Collaboration



Simon
Candelaresi

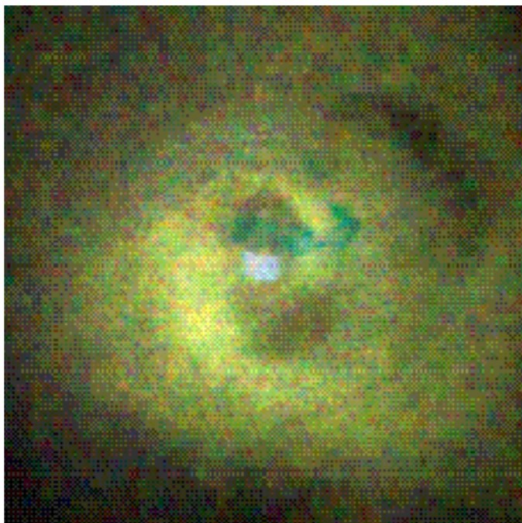
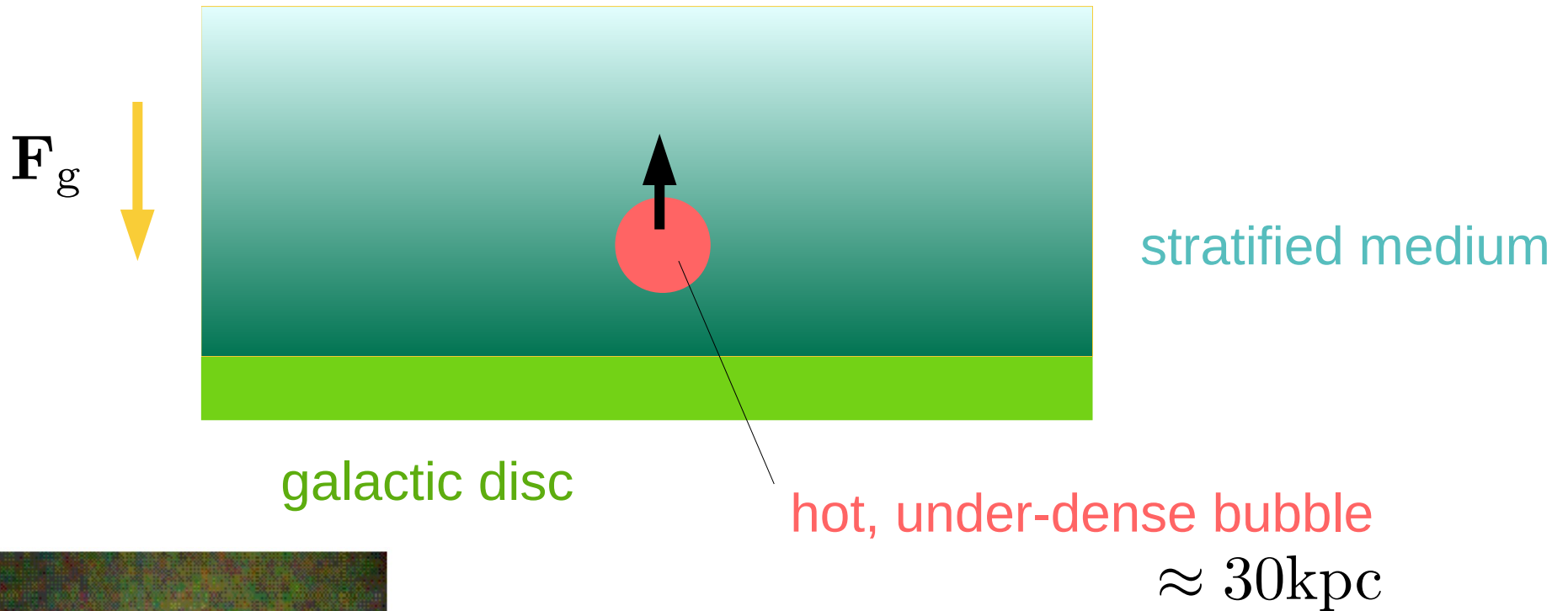
Dundee



Heraklion

Fabio Del Sordo

Intergalactic Bubbles

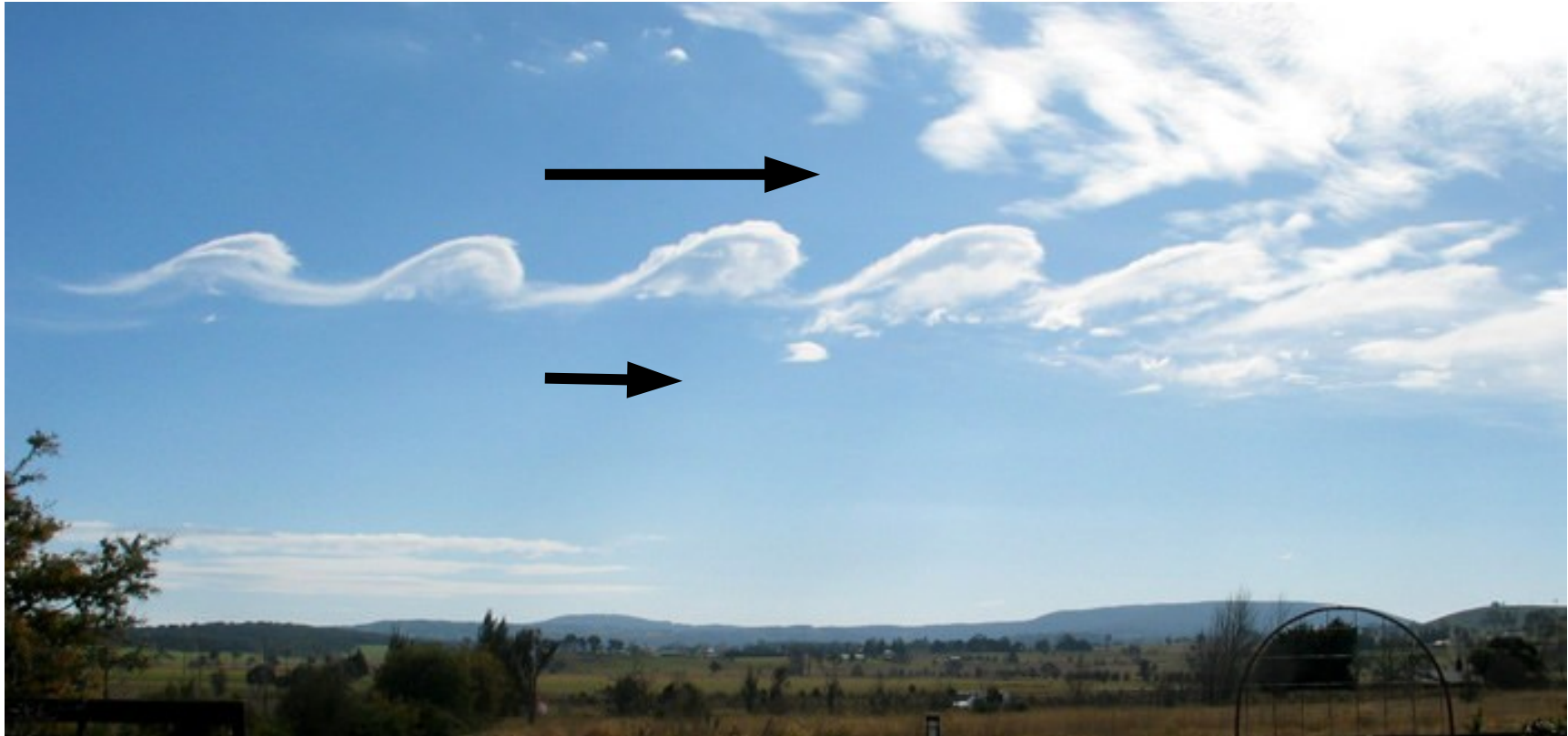


➔ Bubbles rise buoyantly through density difference.

➔ Bubbles' age is several tens of millions of years.

(Fabian et al. 2000)

Kelvin-Helmholtz Instability



(GRAHAMUK/Wikimedia Commons)

➡ Bubbles should get disrupted.

➡ What is the reason for their stability?

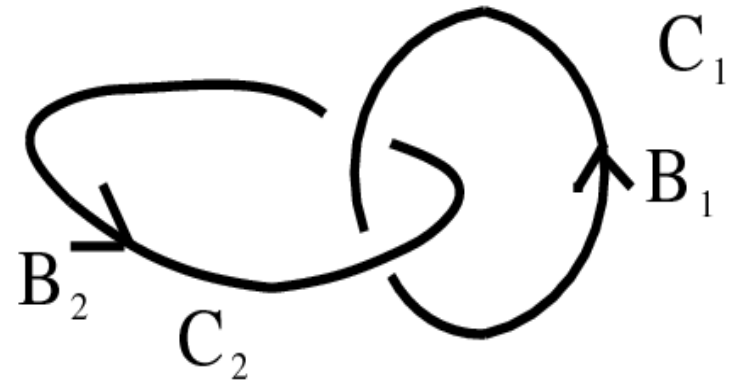
Magnetic Helicity

Measure for the topology:

$$H_M = \int_V \mathbf{A} \cdot \mathbf{B} \, dV = 2n\phi_1\phi_2$$

$$\nabla \times \mathbf{A} = \mathbf{B} \quad \phi_i = \int_{S_i} \mathbf{B} \cdot d\mathbf{S}$$

n = number of mutual linking



(Moffatt 1969)

Conservation of magnetic helicity:

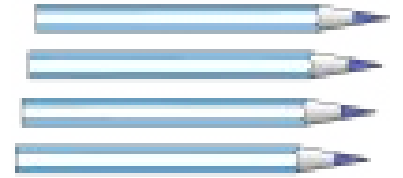
$$\lim_{\eta \rightarrow 0} \frac{\partial}{\partial t} \langle \mathbf{A} \cdot \mathbf{B} \rangle = 0 \quad \eta = \text{magnetic resistivity}$$



Can a helical magnetic field stabilize the bubbles?

Numerical Experiments

Full resistive magnetohydrodynamics simulations with the PencilCode.



$$\frac{\partial \mathbf{A}}{\partial t} = \mathbf{U} \times \mathbf{B} + \eta \nabla^2 \mathbf{A}$$

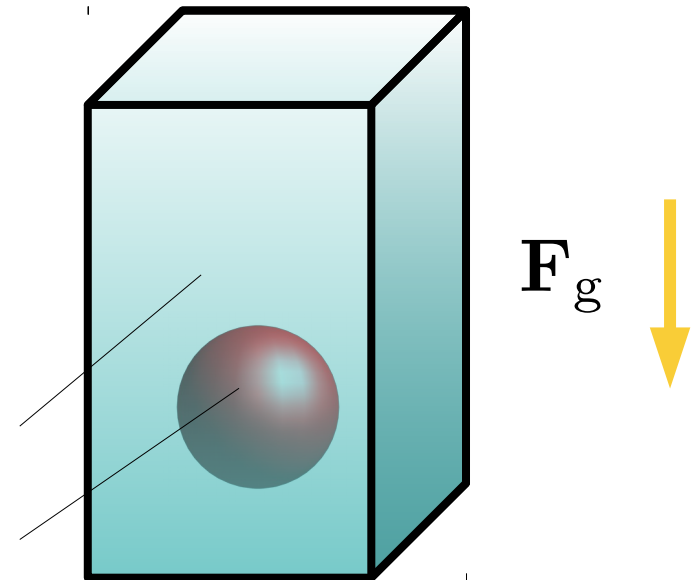
$$\frac{D\mathbf{U}}{Dt} = -c_S^2 \nabla \left(\frac{\ln T}{\gamma} \ln \rho \right) + \mathbf{J} \times \mathbf{B} / \rho - \mathbf{g} + \mathbf{F}_{\text{visc}}$$

$$\begin{aligned} \frac{\partial \ln T}{\partial t} = & -\mathbf{U} \cdot \nabla \ln T - (\gamma - 1) \nabla \cdot \mathbf{U} \\ & + \frac{1}{\rho c_V T} (\nabla \cdot (K \nabla T) + \eta \mathbf{J}^2 \\ & + 2\rho \nu \mathbf{S} \otimes \mathbf{S} + \zeta \rho (\nabla \cdot \mathbf{U})^2) \end{aligned}$$

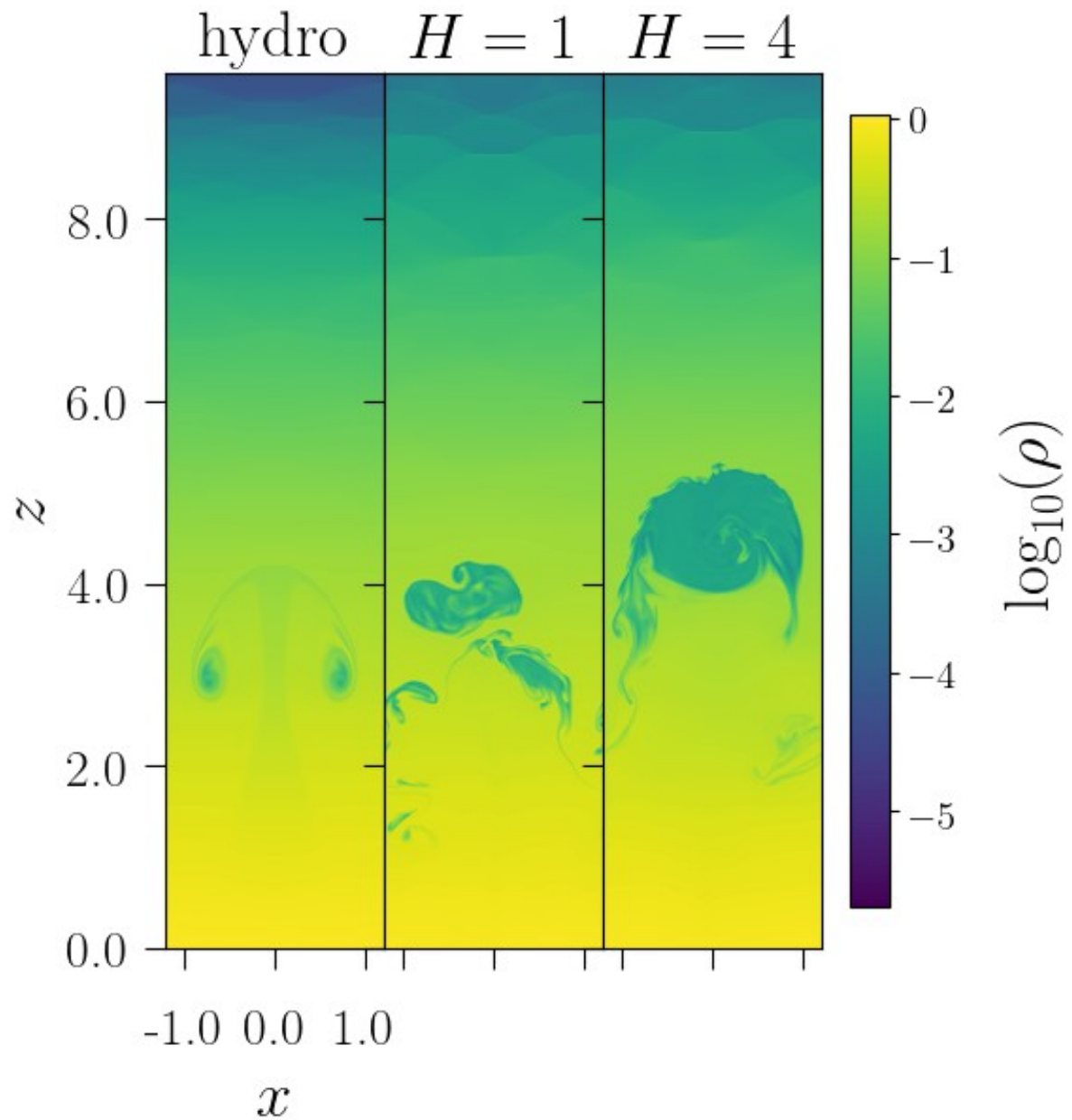
$$\frac{D \ln \rho}{Dt} = -\nabla \cdot \mathbf{U}$$

stratified medium

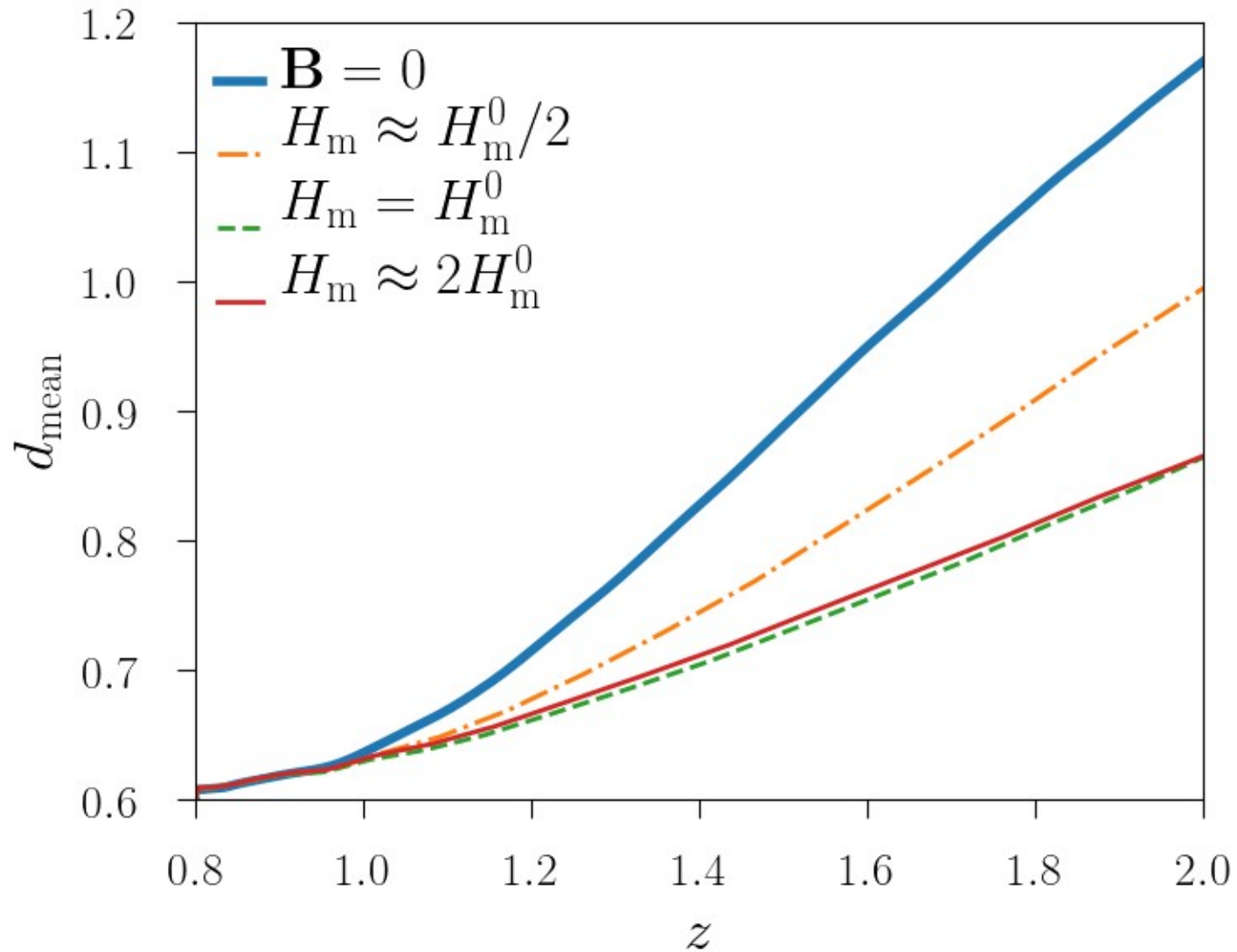
hot, under-dense bubble



Bubble Coherence



Bubble Coherence



Helical magnetic fields can stabilize the bubbles.

Conclusions

- Intergalactic bubbles are very stable.
- Magnetic helicity leads to stability at small magnetic energy.
- Combined expertise between Heraklion and Dundee.